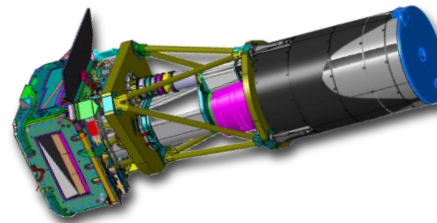
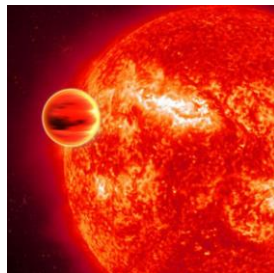




CoRoT-7 b

The discovery of the
first transiting Earth-like
planet by the CoRoT satellite



Manne Siegbahn Memorial Lecture 2010 – Stockholm

D. Rouan, Observatoire de Paris



1- Exoplanets hunting

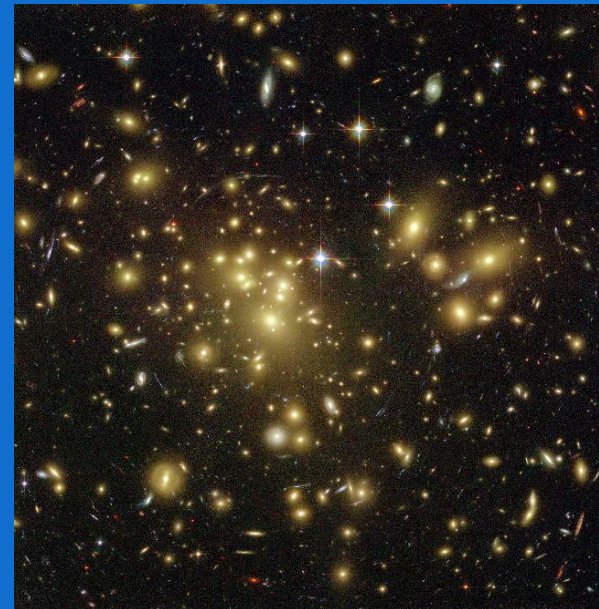


Other planetary systems ?

- The sun : a star among 100 billions in the Milky Way



- The Milky Way : one galaxy among 100 billions in the *accessible* Universe
- The sun would be unique among 10^{22} stars ?



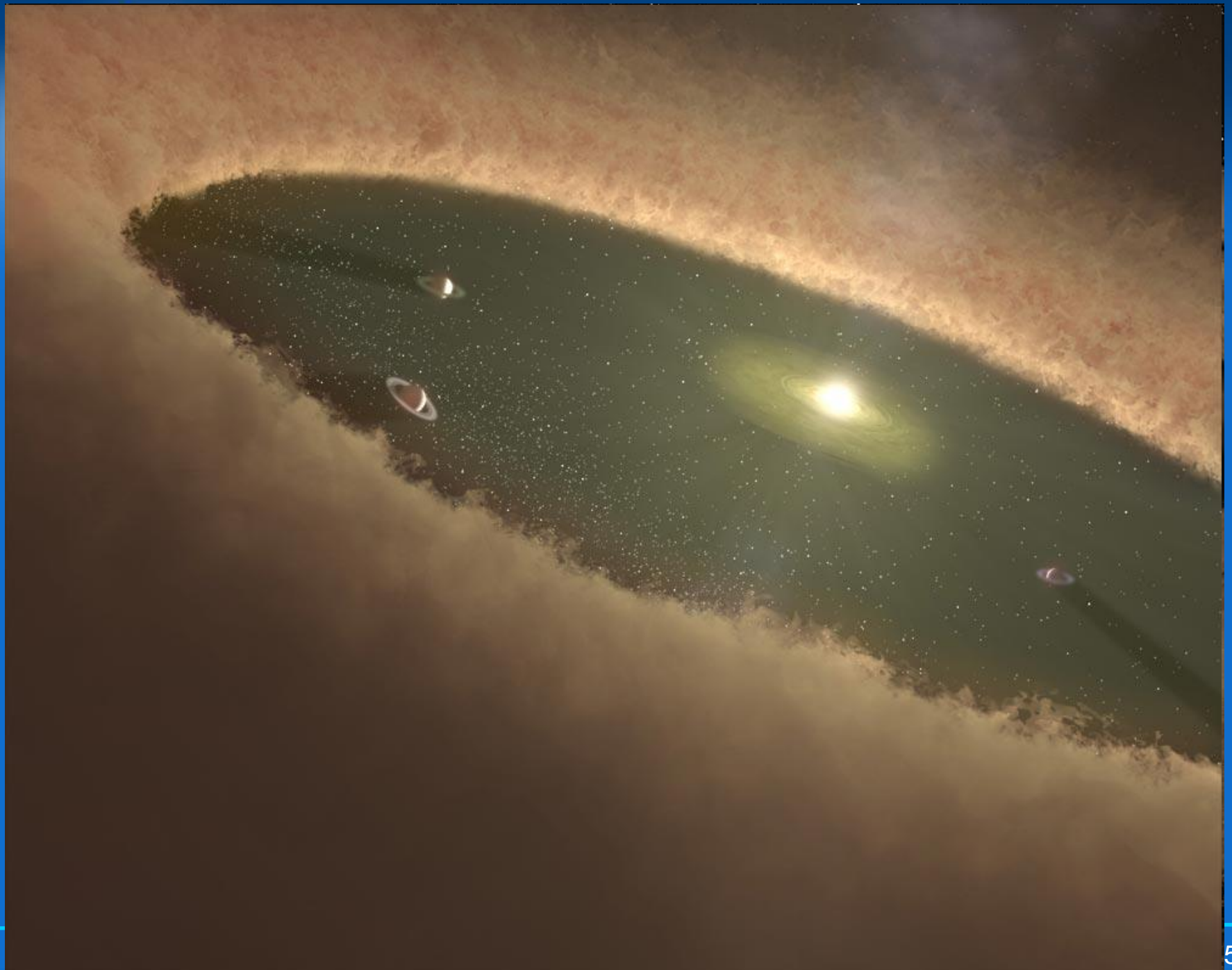


Other planetary systems ?

- Astronomers are convinced since a long time that the **solar system** with its 8 planets and its « small bodies » (asteroids, moons, satellites), **is not unique in the Universe**.



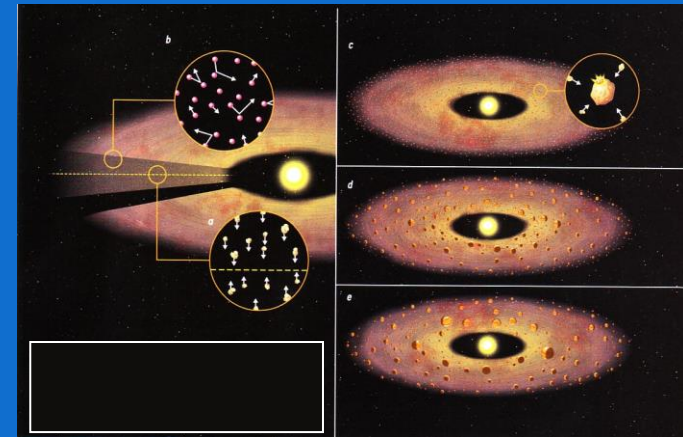
- However, it's only since 1995 that they got a proof, with the established discovery of **the first extrasolar planet** by M. Mayor & D. Queloz : **51 Peg B**





Why detecting other planetary systems ?

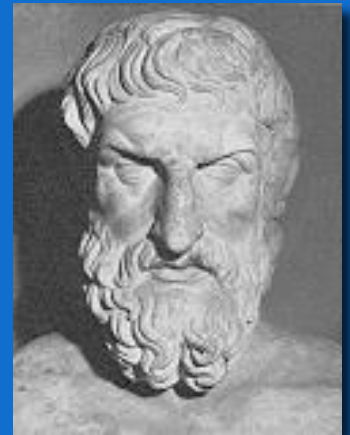
- To understand in details how the star/planets system forms :
 - under which necessary conditions ?
 - what is responsible of the observed broad variety ?
- To predict the evolution of our own system and of each of its individual planets
 - Migration of planets from outside to the interior
 - Planets *ejected* by tidal interaction with other planets
 - Planets loosing their atmosphère (Mars) or becoming hot (Venus)





Why detecting other planetary systems ?

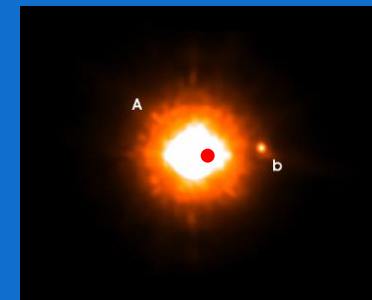
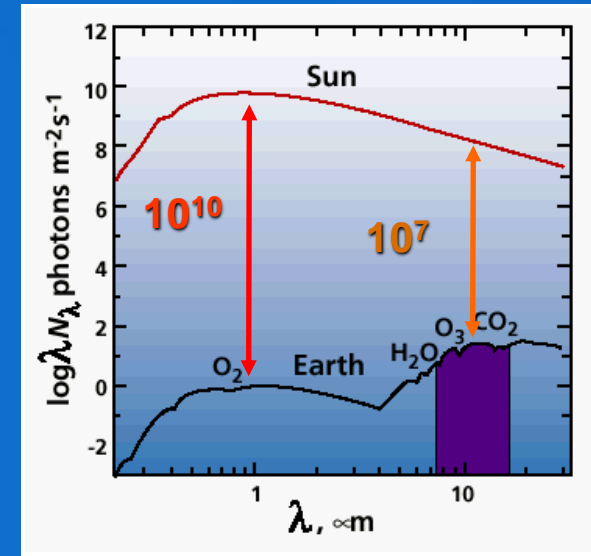
- Detecting other islands of life in the Universe ?
- An ancient and fundamental question
 - « Other worlds, with plants and other living being, some similar and other different from ours must exist »
(Epicurus, 300 B.C)
- A question accessible to anyone, that has resonance in any of us
- A question to which science pretends to bring a first element of answer in a rather near future





Detecting directly a planet ?

- An exoplanet is difficult to detect directly because :
 - It is much smaller than the star it is orbiting
 - It is not self-luminous : it just reflects the light of the star.
- It's much less bright than the star
- It's very close from the star that dazzles the observer



Detecting directly a planet ?

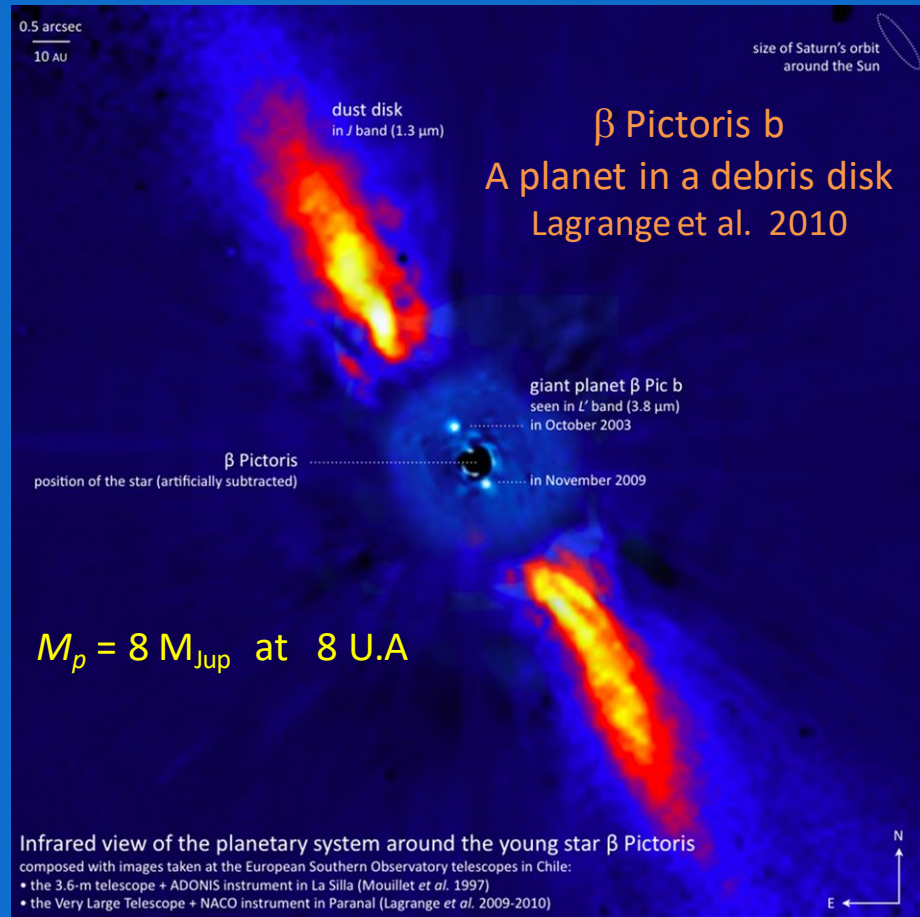
*Direct detection of a
planet is
AWFULLY difficult !!!
Thus indirect methods
have been the most
productive until now*



Detecting directly a planet !

- Indirect detection is the rule for the 490 exoplanets that have been found today
- However, in very few cases corresponding to especially favourable situations, **direct detection has been possible** :

Beta Pictoris B : a planet on a 8 AU orbit (\approx Saturn) around a young star.

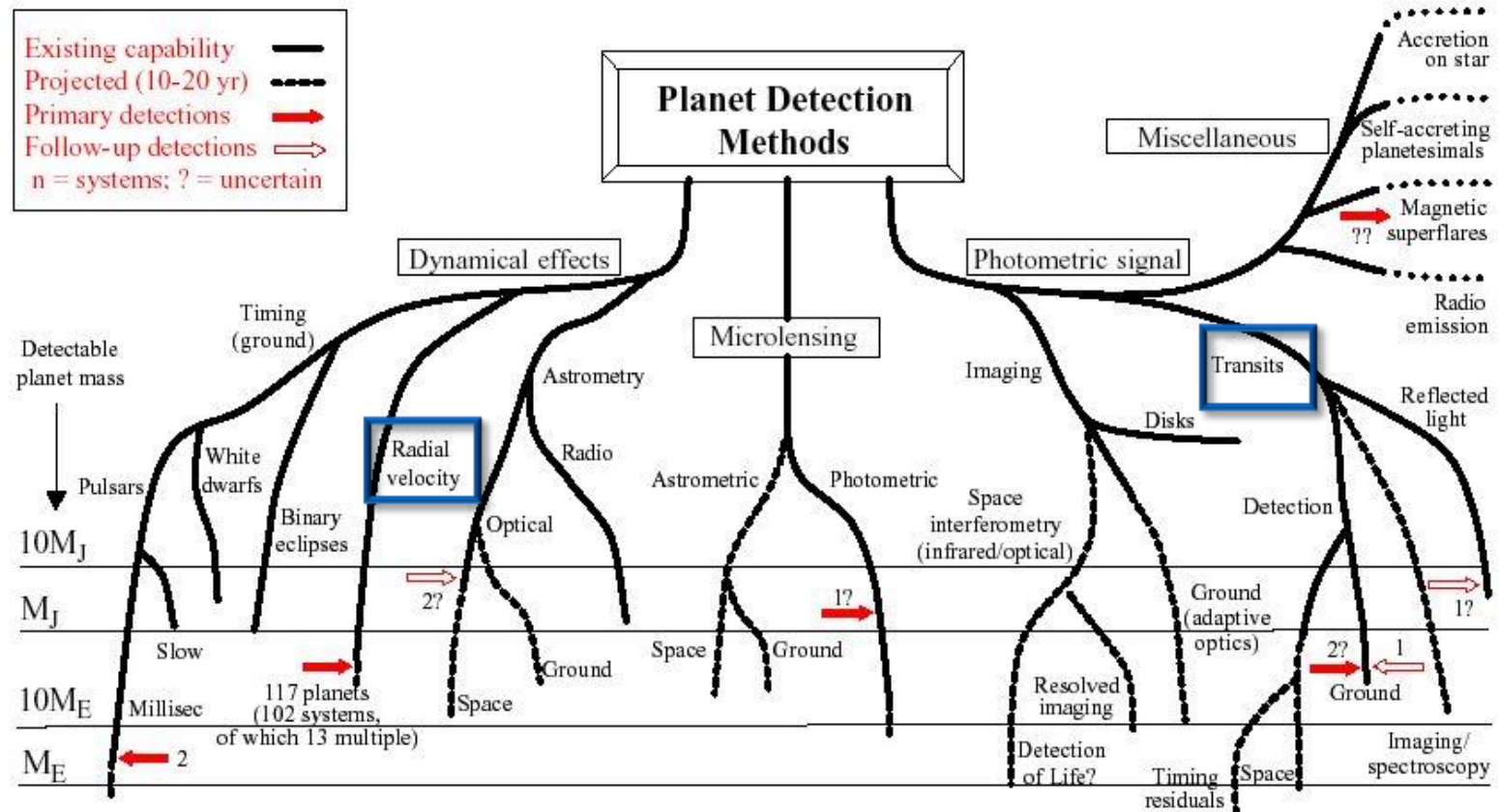




The tree of planet detection methods

Planet Detection Methods

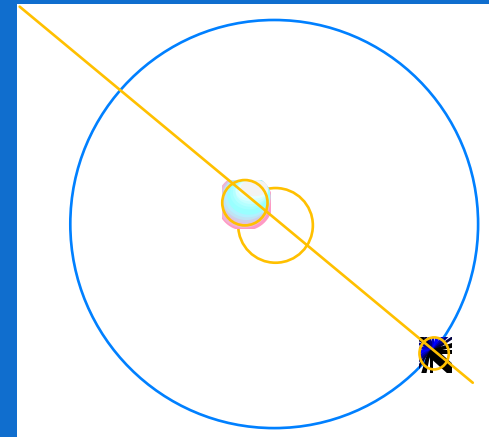
Michael Perryman, Rep. Prog. Phys., 2000, 63, 1209 (updated September 2003)



The reflex motion



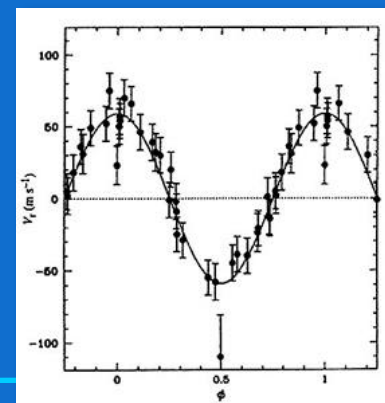
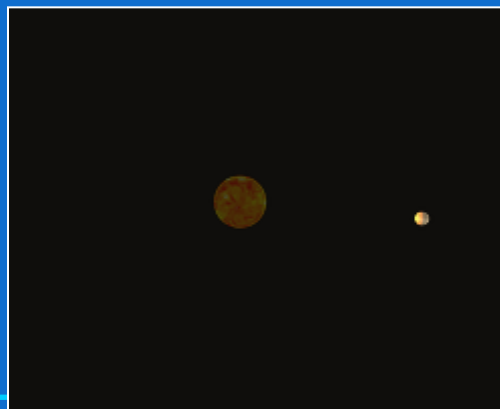
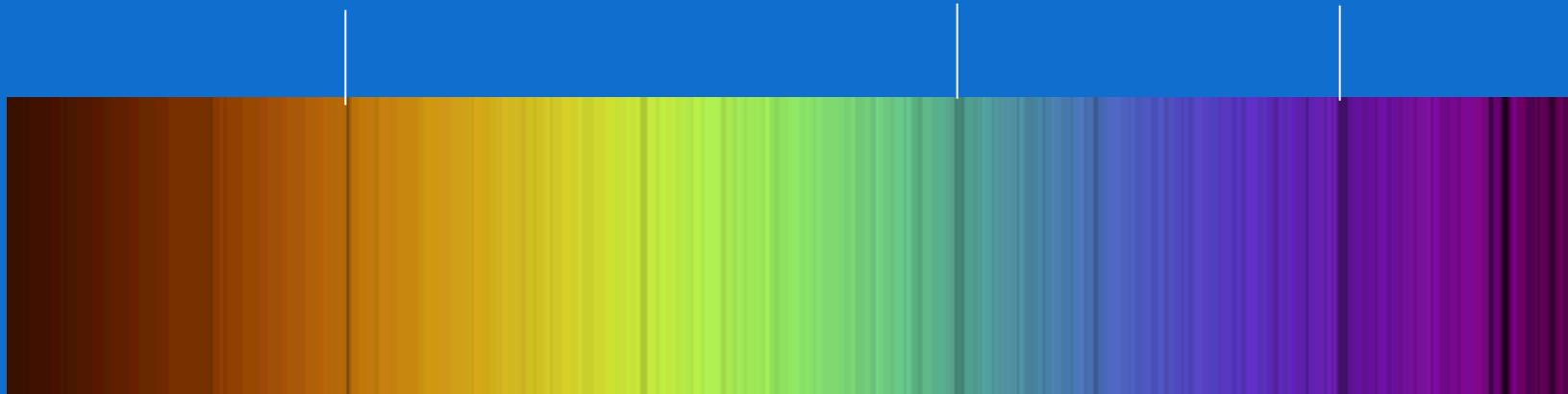
- The center of mass of the star-planet system being fixed, the star travels also on a circular orbit, of very small radius.
- The **radial velocity** of the star on this orbit is measurable thank to **Doppler effect**





Radial velocity measurements

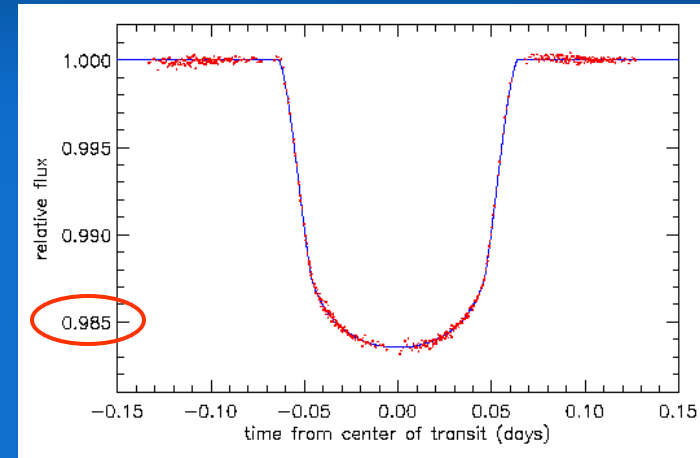
- Using a high resolution spectrograph, one measures the tiny change in wavelength on a large number of spectral lines





Detection of transits

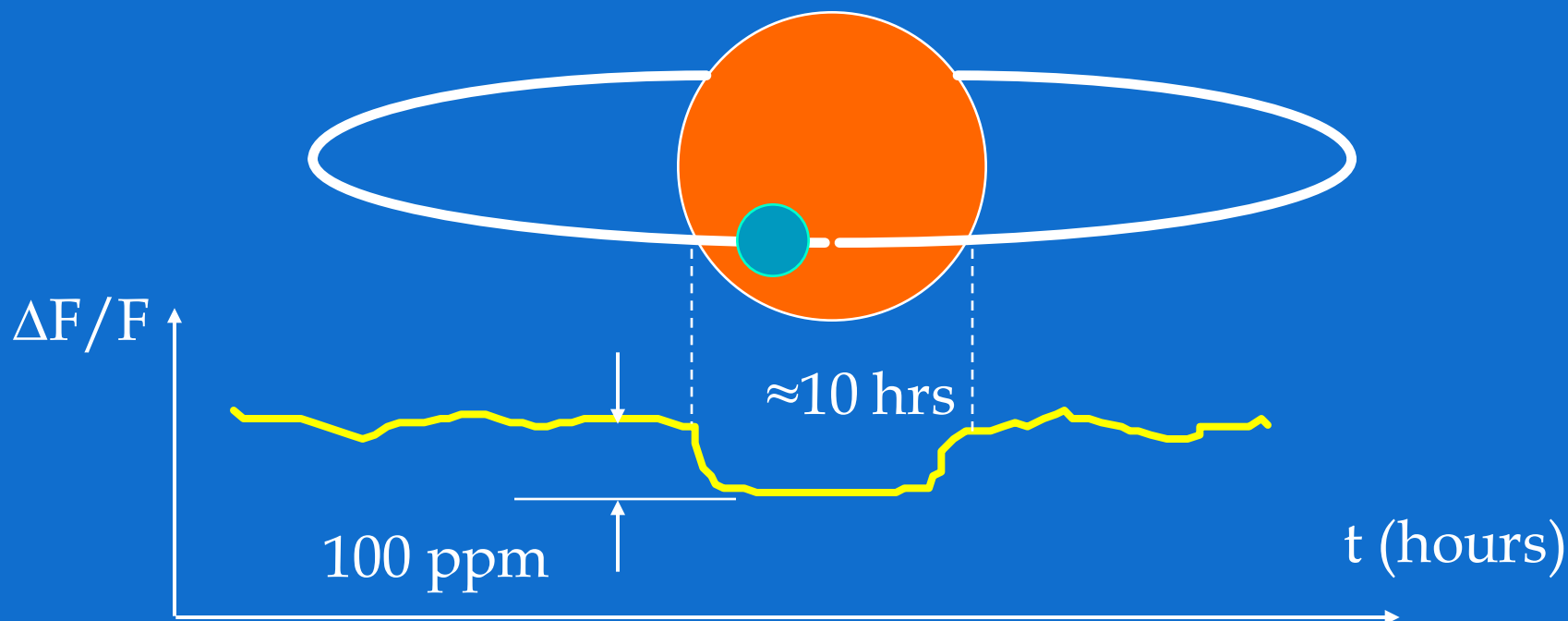
- One looks for events (transits) where the planet travels *between the star* and the observer → small decrease of star brightness





Transit of a planet

- The chance to be in favourable case (observer in the orbital plane) = 0.5% for the Earth vs Sun
- Consequence: one must monitor a large number of stars during long periods





Transit : few basic relations

- Signal $\approx (R_p/R_*)^2$ (+ limb darkening effect)
Jupiter / Soleil = 1% ; Earth/Sun : 0.01 %

- Probability for the observer to be in the orbital plane

$$p = R_*/a$$

Earth / Sun $p = 0.5\%$

51PegB $p = 10\%$

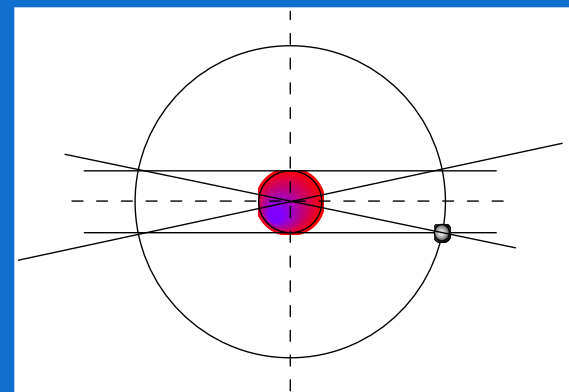
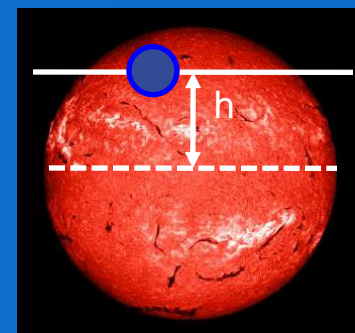
- Transit duration :

For an impact parameter h

$$\tau = 14 \text{ h } a^{1/2} M_*^{-1/2} R_* (1-h^2)^{1/2}$$

Earth: 14h ; Jupiter : 31h ; 51PegB : 3h

- Cumulated Signal : $N \tau \propto a^{-1}$: Favour planets w small orbit





Detection of transits : a specially rich information



- **Mesure of planet's size**

If measure of the mass \Leftrightarrow **density** and thus clues on the structure

- Impact parameter \Leftrightarrow Inclination of the orbit
- Secondary transit \Leftrightarrow albedo and temperature
- Photometric variations during the transit \Leftrightarrow stellar spots
- Rössiter effect \Leftrightarrow rotation axis of the star vs orbital plane
- Variations in periodicity \Leftrightarrow other planets or satellites
- Spectroscopy of transits : absorption or emission lines \Leftrightarrow composition



The CoRoT satellite



COROT : a European project

- *CoRoT : Convection Rotation Transit*
- France : CNES main contractor
 - Alcatel Alenia Space
 - Institutes : LESIA (Observatoire de Paris), LAM (Marseille), IAS (Orsay)
- Other partners : Austria, Spain, Germany, Belgium, ESA, Brazil
- Launched on 27 December 2006
- Mission recently extended for 2 more years



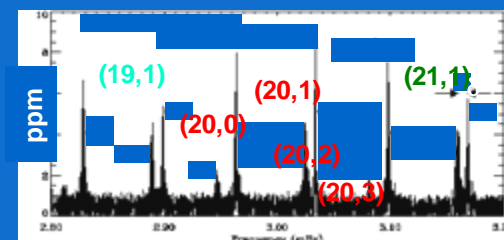


CoRoT: a photometric satellite

- Principle: monitor continuously the flux of thousands of stars during 10 cycles of 5 months each
- Two programmes conducted in parallel :

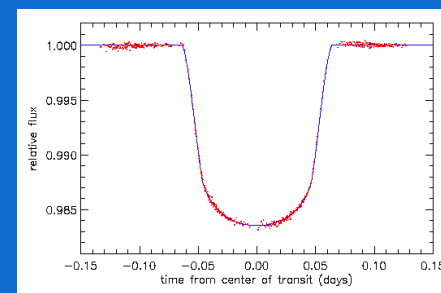
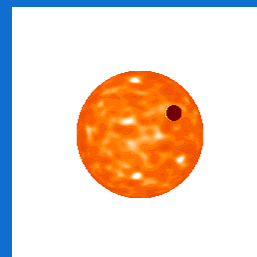
- **Asterosismology:**

- ~ 15 étoiles
- One exposure / sec



- **Exoplanets (transits) :**

- ~ 12.000 stars
- One exposure /32 sec
- 3 colors





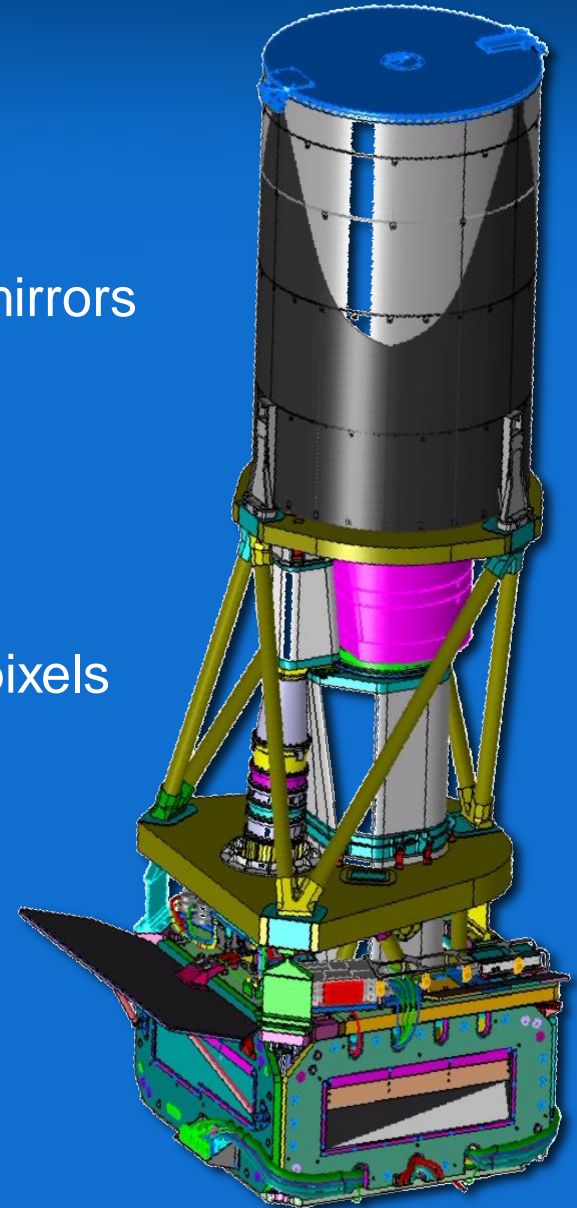
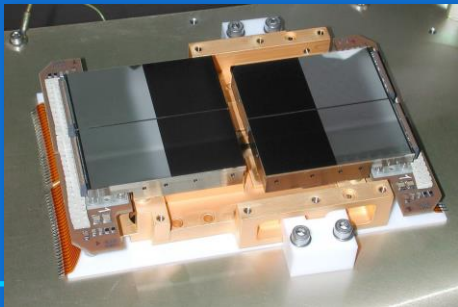
The CoRoT Instrument

● Télescope

- Pupil diameter : 27 cm
- Afocal telescope : 2 parabolic mirrors
- Long sunscreen

● Wide field camera

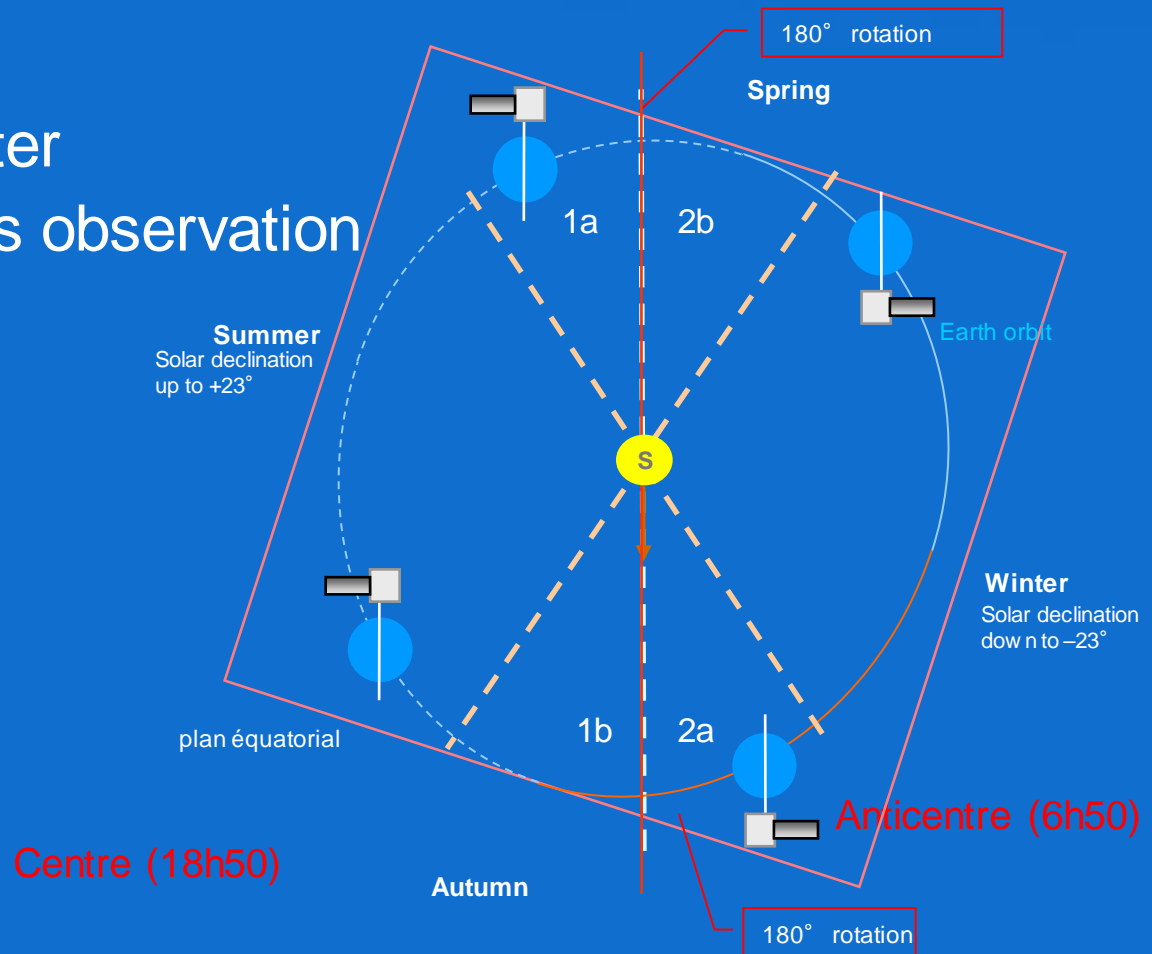
- Dioptric objective w 6 lenses
- Focal plane : 4 CCD of 2048^2 pixels
 - 2 CCD astero-sismology
 - 2 CCD exoplanets
- Exoplanet field : $3.45(^{\circ})^2$





Observing strategy

- Polar orbit
- Sun *in the back* →
Flip each semester
- 150 days of continuous observation of each field
- 10 fields

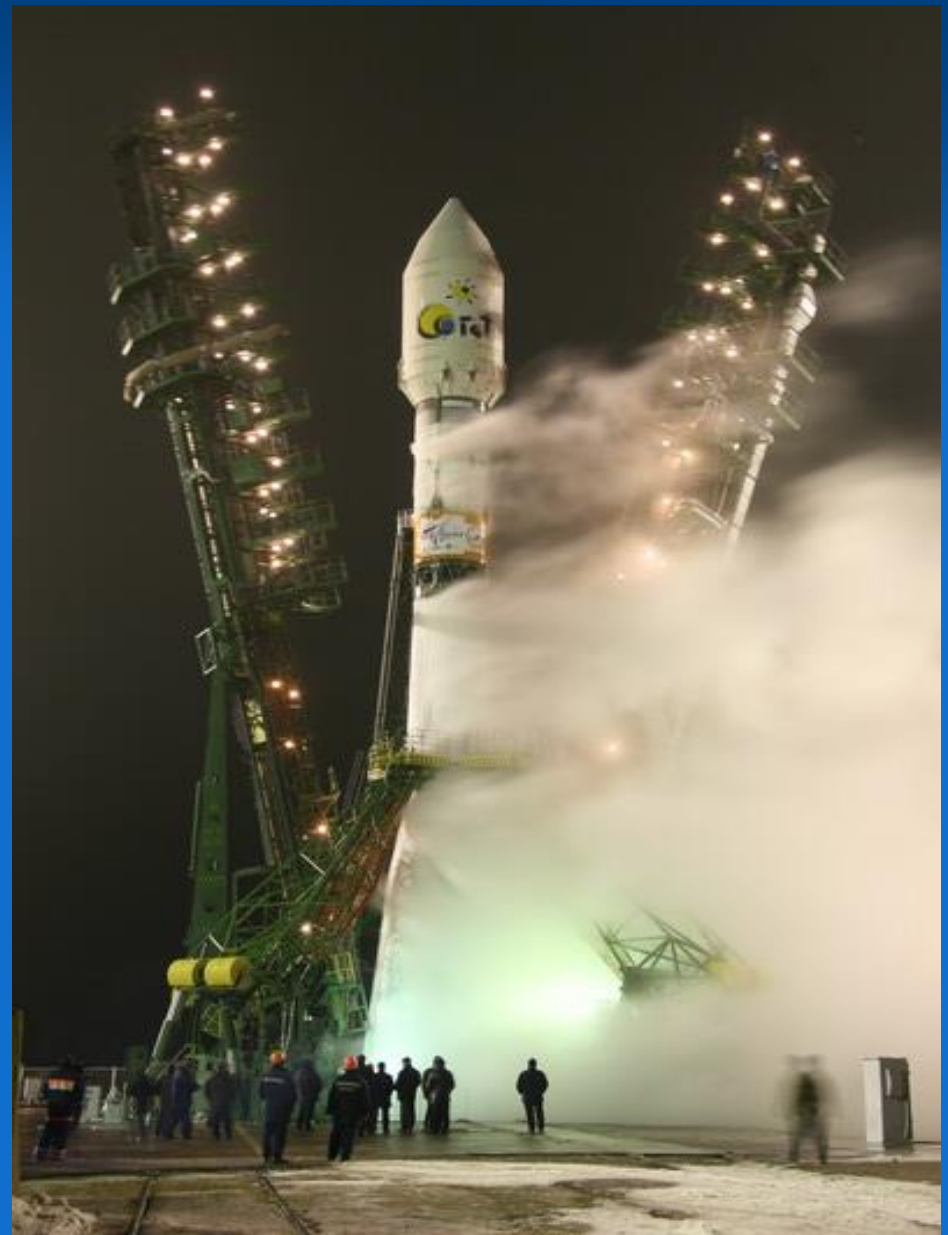




Launch

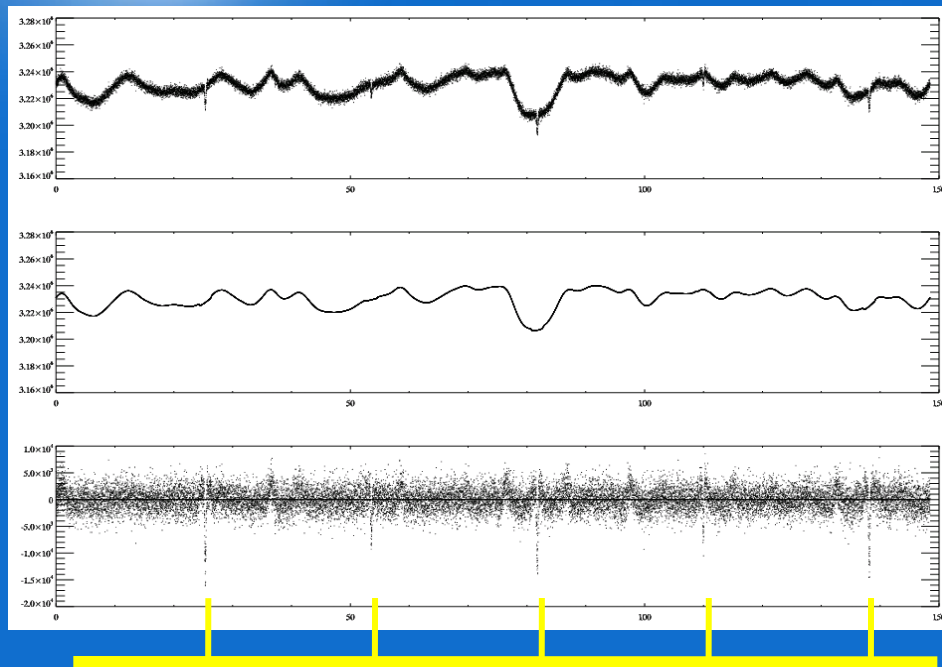


With a good old faithful Soyuz





Processing a light curve



Original

Low frequencies :
stellar fluctuations

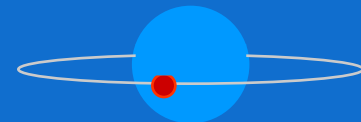
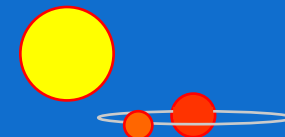
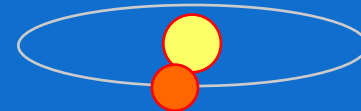
High frequencies :
transits

- Detection of a planet candidate : at least **three** transit-like periodic events
→ orbital periods $T < 50$ days



Beware of *false positives*

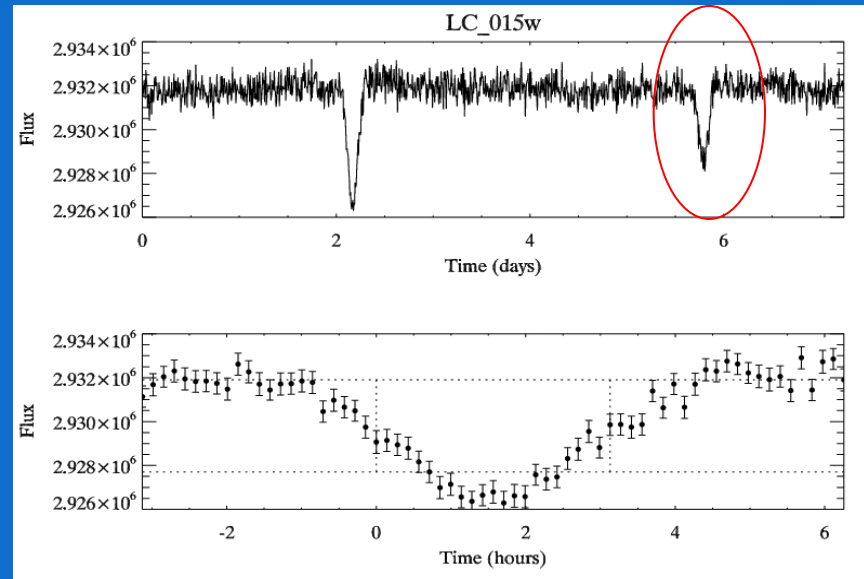
- The enemy : the **false positive** !
produces a signal that mimics a transiting planet
- The star is a Grazing Eclipsing Binary (GEB)
- Eclipsing Binary in the Background (BEB)
- Eclipsing Binary in a dwarf/giant system
- FOLLOW-UP FROM THE GROUND IS MANDATORY !





How identifying a liar ?

- Transit : V-shape and secondary eclipse

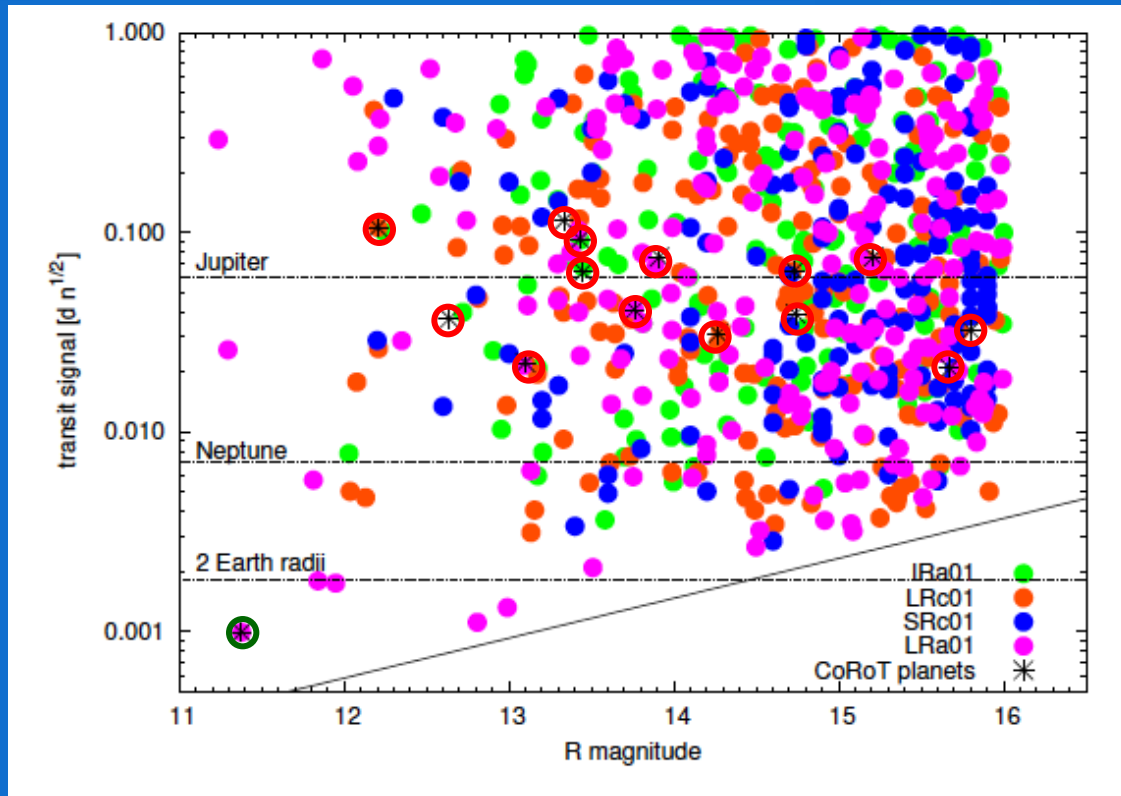


- Radial Velocity → mass of star or of a planet
- Spectroscopy → identify giant stars or blend (binary)
- ON/OFF Photometry of nearby stars → background binaries
- CoRoT Colours of the transit → binary system



CoRoT does detect planets !

- 17 planets up to now (14 published)



- among which one of special interest ...

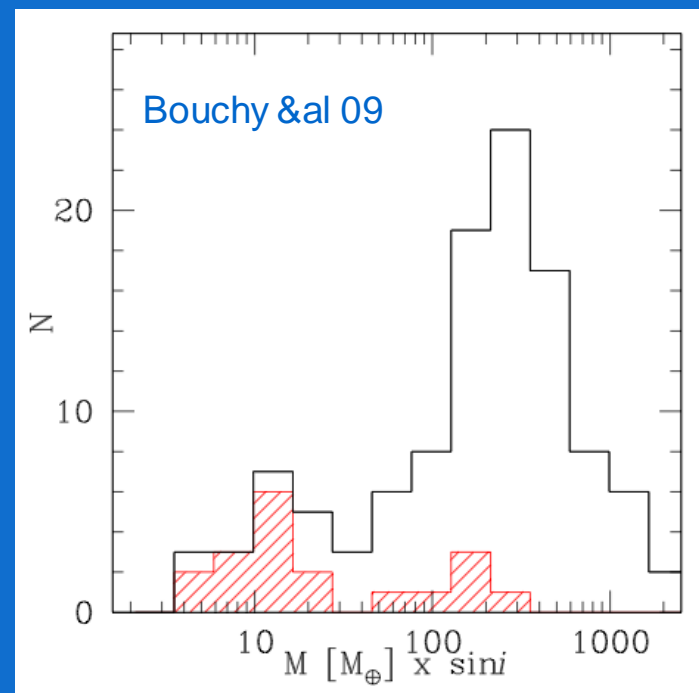


CoRoT-7b: an Earth's cousin



Short period planets:

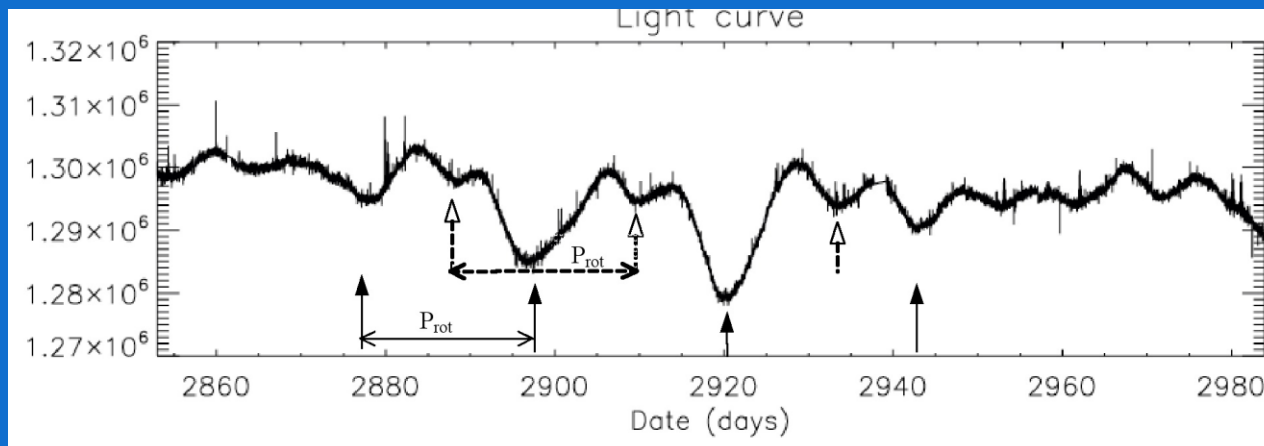
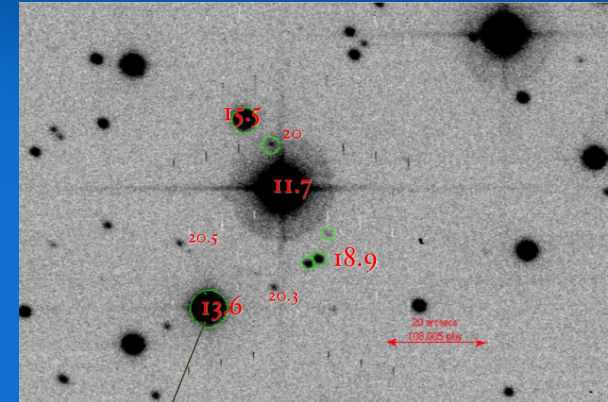
- 2006: first Super Earth detected (gravitational lensing)
 - Beaulieu et al. Nature 06
- Recent results from Radial Velocity point to a very significant population of Super-Earth: e.g. Mayor et al. 09
- Today: 55 planet w $M < 0.1 M_{\text{Jup}}$
 - An important component: 30 % of all solar type stars may harbour one hot SE
- Size needed !
 - to derive the structure
- One major goal of **CoRoT**
 - detect transits of Super-Earth
 - performances allow it
 - And indeed a first case !





The star CoRoT-7

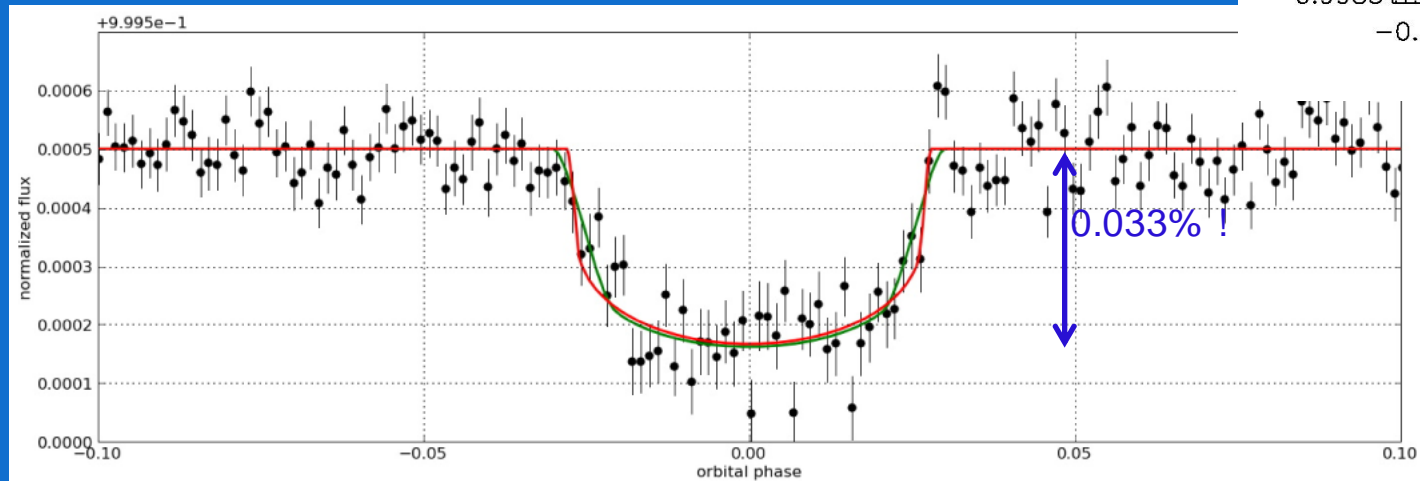
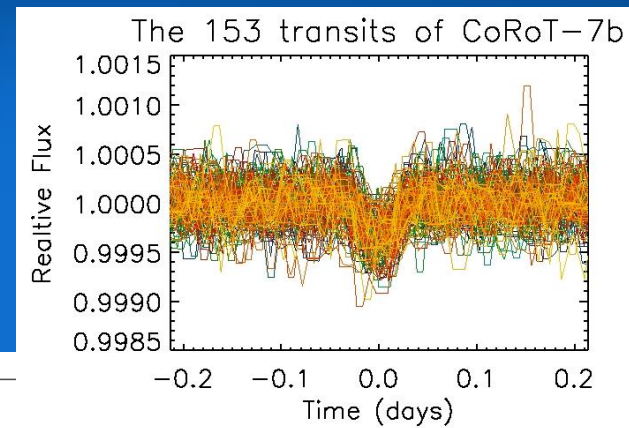
- Spectrum → type
- Parameters :
 - G9V star - $m_V = 11.7$
 - $T_{\text{eff}} = 5250 \text{ K}$: quasi-solar
 - Distance = $130 \text{ pc} \pm 30$
 - age $\approx 1.2 \text{ Gyr}$
- Very active : emission feature in H and K Ca lines, 2% variability !



- Spots clearly seen crossing the disk because of rotation
→ rotation period = 23 days



The folded light curve

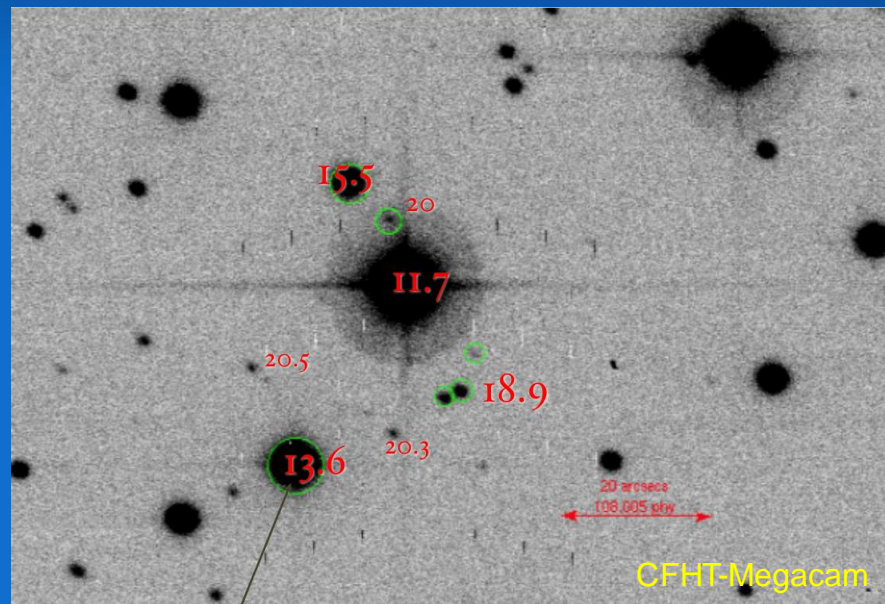


- Transits first detected by quick-look analysis (« alarm » mode)
- 153 transits, all ~ seen when superimposed
- Extremely short period : $P = 0.8536$ days
→ very small transit depth : $\Delta F/F = 0.033\%$



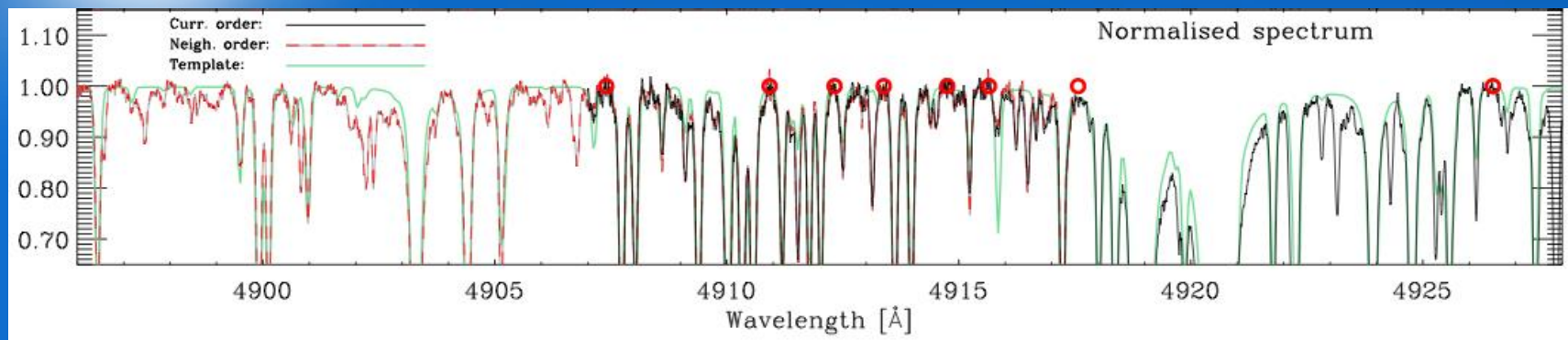
A vigorous follow-up programme

- Such a small transit signal
- IF A PLANETARY ONE -
means a very small planet :
 $R_{pl} = 1.7 R_{earth}$
- A small planet or a usual case of false positive, such as BEB, GEB or giant planet in a triple system ?
- → A vigorous Follow-up programme from the ground, using several of the top-level instruments in the world was conducted

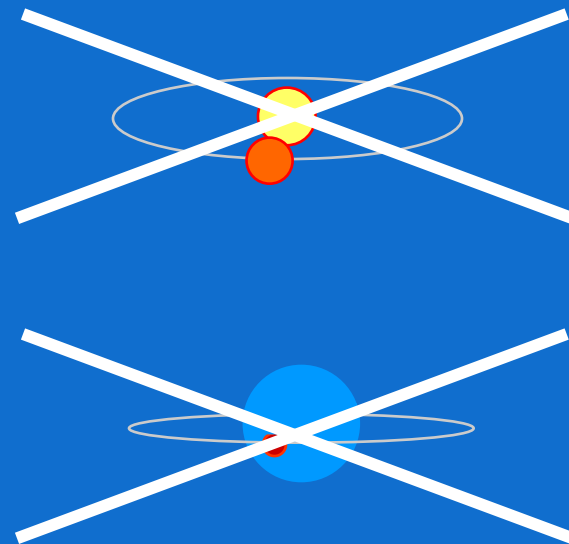




Follow-up 1 : Spectrum



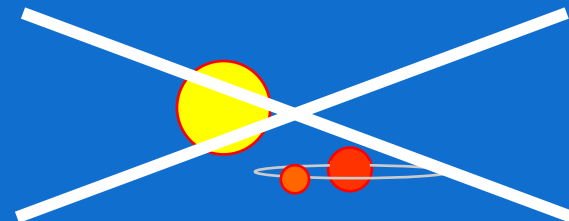
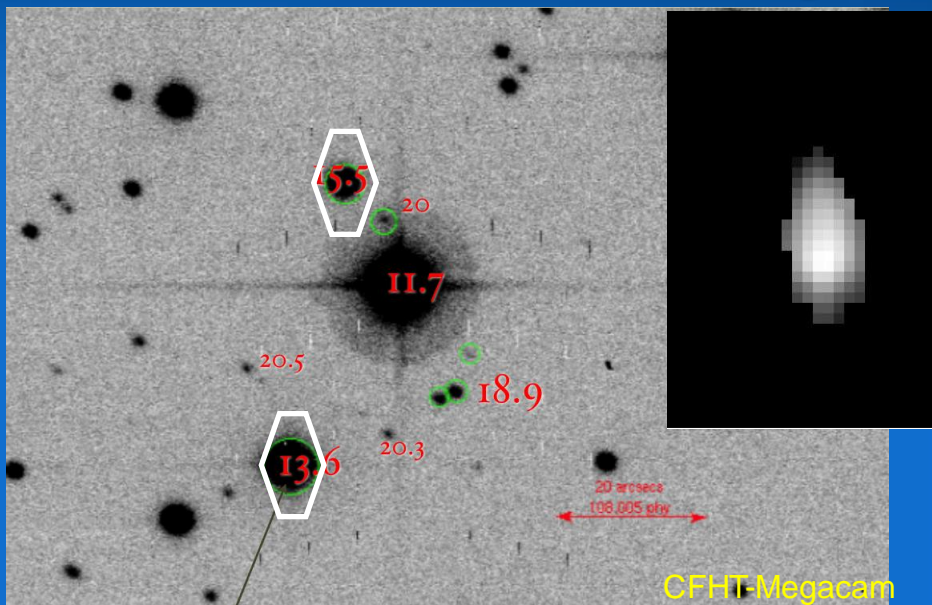
- High resolution spectra
- → a unique star of type G9V
- Cannot be a blend
- Excludes a Grazing Eclipsing Binary
- Excludes a giant star eclipsed by a dwarf star or a giant planet





Follow-up 2 : ON/OFF Photometry

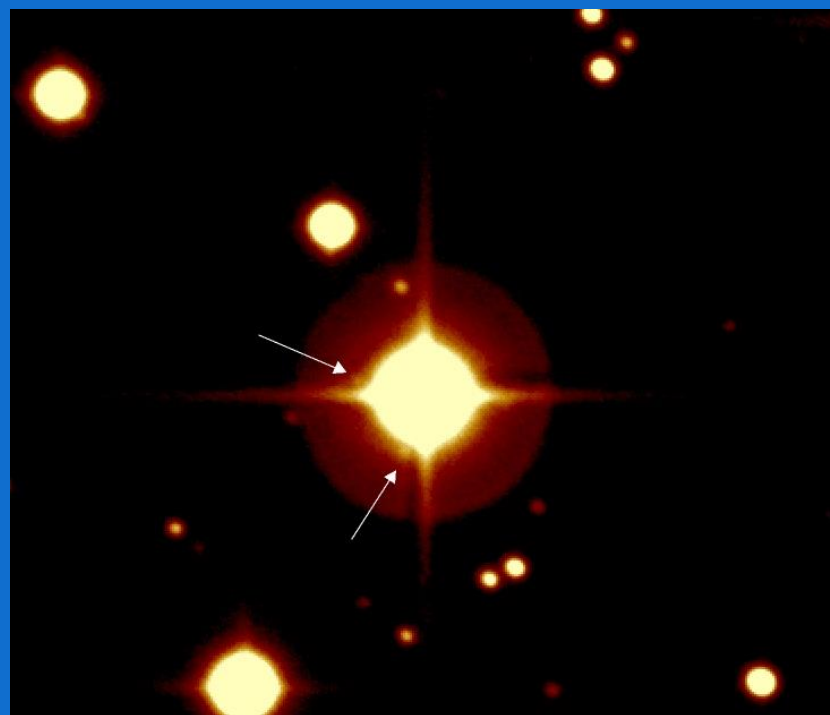
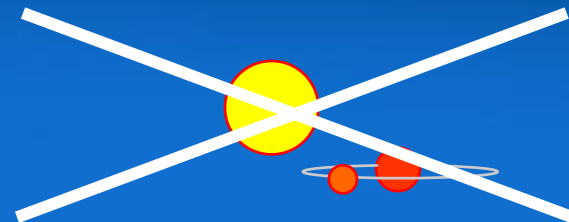
- ON/OFF CCD photometry on nearby stars
- PSF + Corot mask \rightarrow light in corot-aperture :
 - 99.63 % from Corot-7b
 - 0.24 % $V=15.5$ * 17" NE
 - $<0.1\%$ $V=13.6$ * 30"SE
- Only those 2 stars could produce false alarm
- ON/OFF photometry :
 - NO $\Delta\text{mag} > \text{alarm}$
- BEB at distance 8 to 30 arcsec excluded





Follow-up 3 : Good seeing Imaging

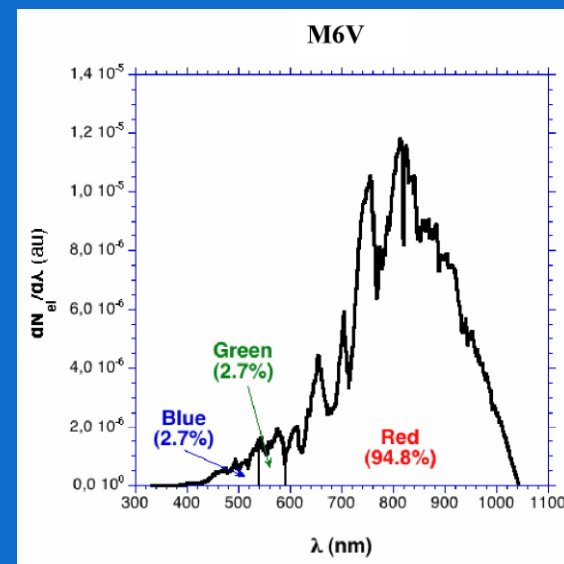
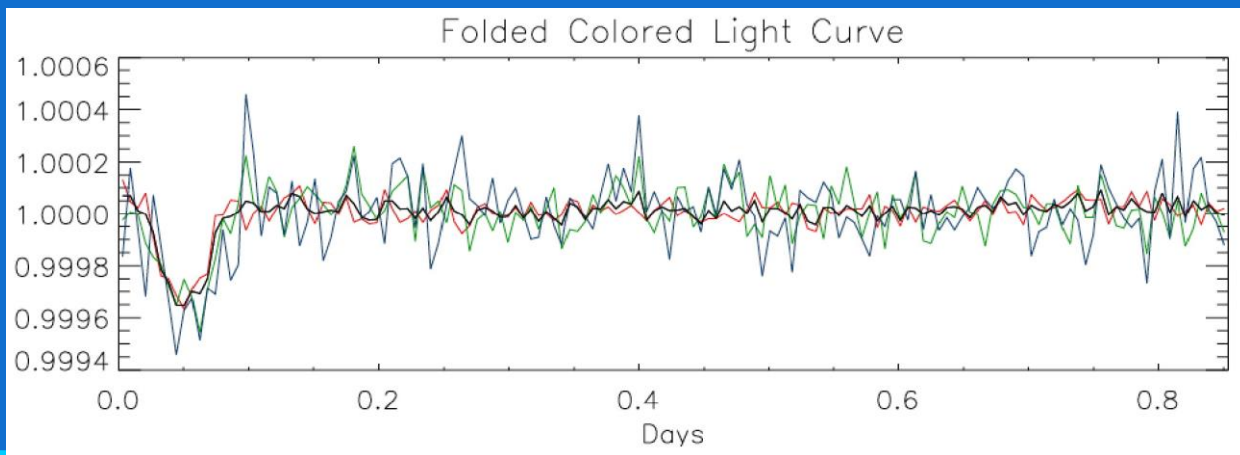
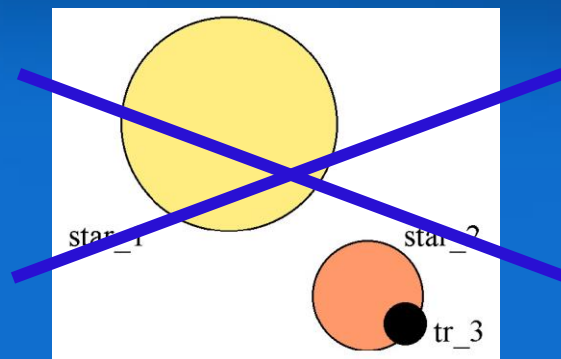
- Good angular resolution images in a very good site (CFHT Hawaii)
 - sub-pixel recentering
 - median of the cube
- Two faint stars detected @ 4.5 & 5.5 arcsec
- Magnitude estimate (fake stars added) : $\Delta m = 10$
 - too faint to cause a $3.5 \cdot 10^{-4} \Delta F/F$
- BEB at distance 4 to 8 arcsec excluded





Follow-up 3 : CoRoT colors

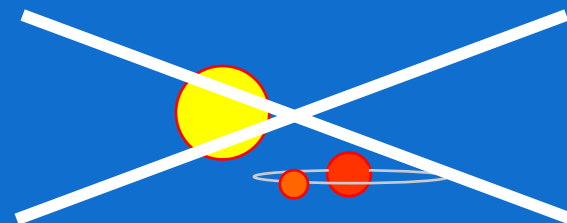
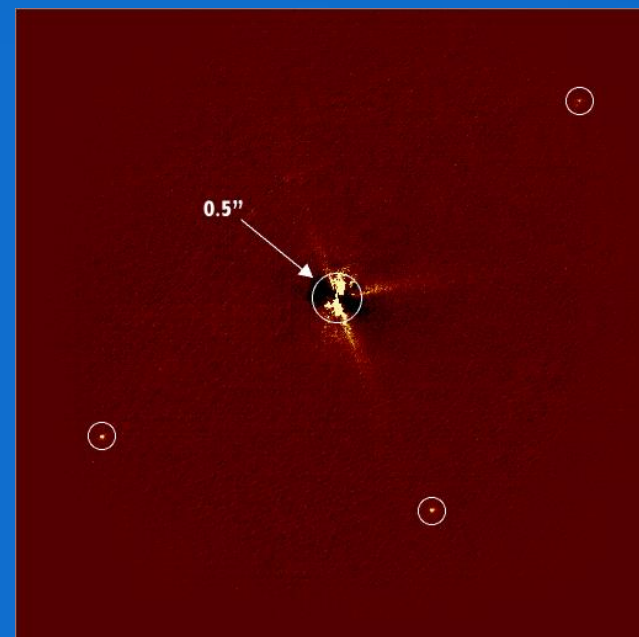
- Triple system with a Jupiter transiting the second star ?
- CoRoT Colors : 3 Light Curves
 - *Blue, Green, Red*
 - transit is achromatic !
- Eliminate a triple system because star 2, and thus the transit signal, should be red





Follow-up 4 : Adaptive Optics Imaging

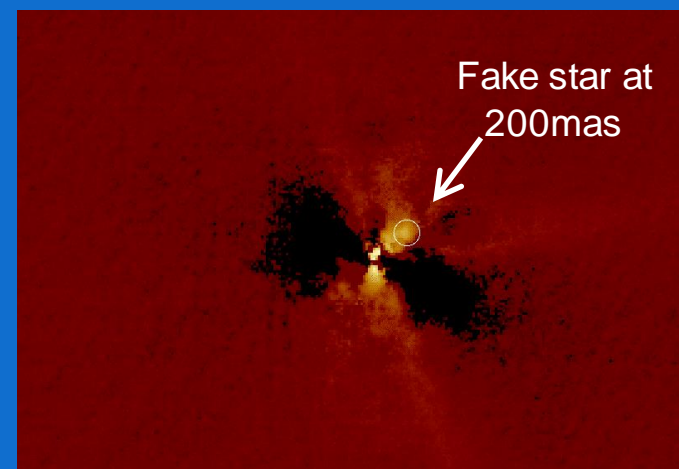
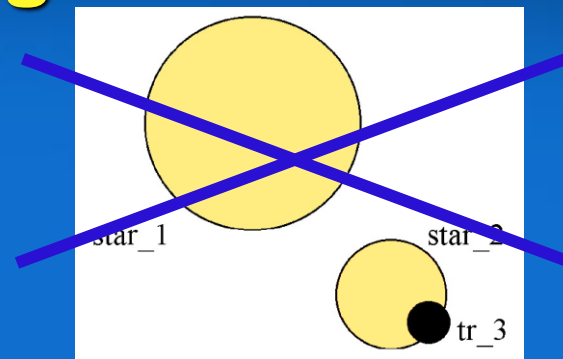
- High angular resolution image in near-IR with NACO-VLT
Adaptive optics + differential imaging by rotation of the camera on the sky
- Three stars detected
all at angular distance $> 4''$
Magnitude = 18.4 - 18.7
too faint to be BEB false positive
- BEB between $0.5''$ to $4''$ excluded





Follow-up 4b: Adaptive Optics Imaging

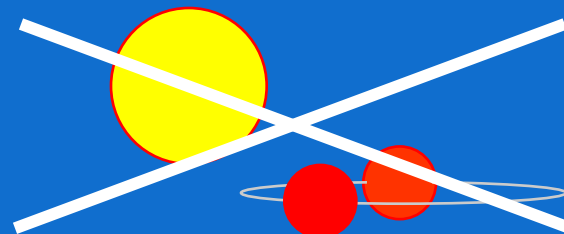
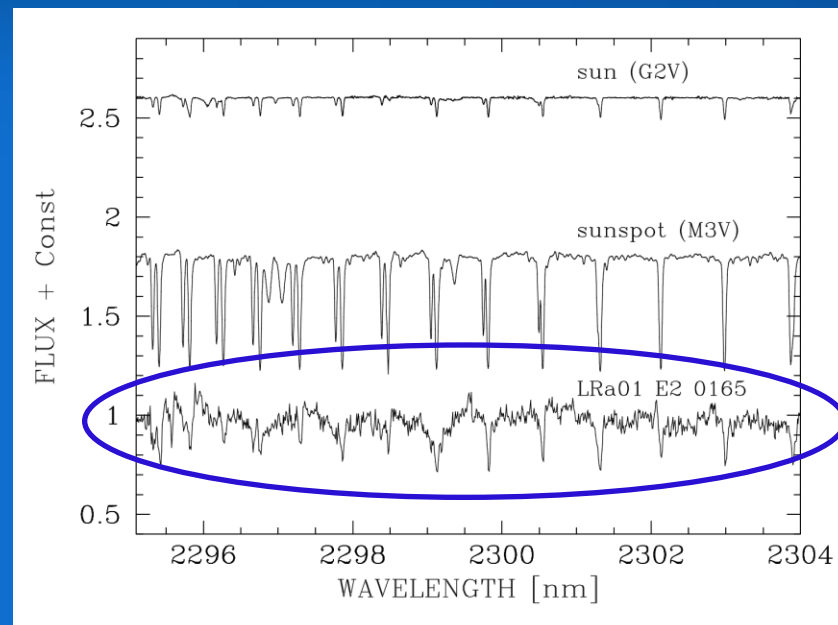
- Jupiter transiting a close background star of same color as Corot-7 ?
 - Would not be detected in Corot colors
 - the star must be $3.5 \cdot 10^{-2}$ fainter
- High angular resolution image with NACO-VLT
 - add a fake faint star at $0.2''$ and $0.3''$
 - do the same processing as before
 - a star with $\Delta m = 5$ would be detected at 300 mas and likely at 200 mas
- Probability to have a star of mag 16.5 and same color, w a transiting Jupiter within $0.25'' < 2 \cdot 10^{-4}$





Follow-up 5 : IR spectra

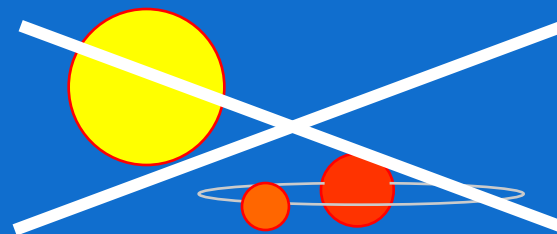
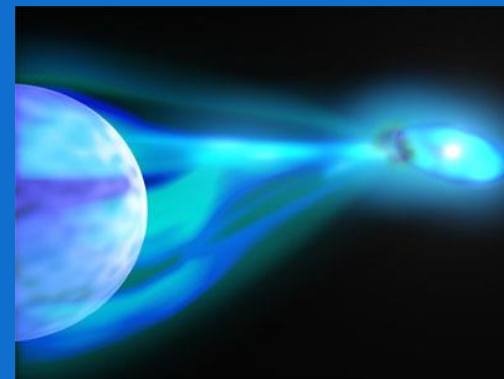
- Triple system with 2 eclipsing low mass stars?
- IR spectra (CRIRES-VLT) :
At 2 μm : strong CO overtone lines in spectra of K and M stars
- Result:
No K or M stars at the distance of Corot-Exo7 within 0.3''
- Eclipsing binary system of M stars orbiting Corot-Exo7b excluded





Follow-up 6 : X-ray activity

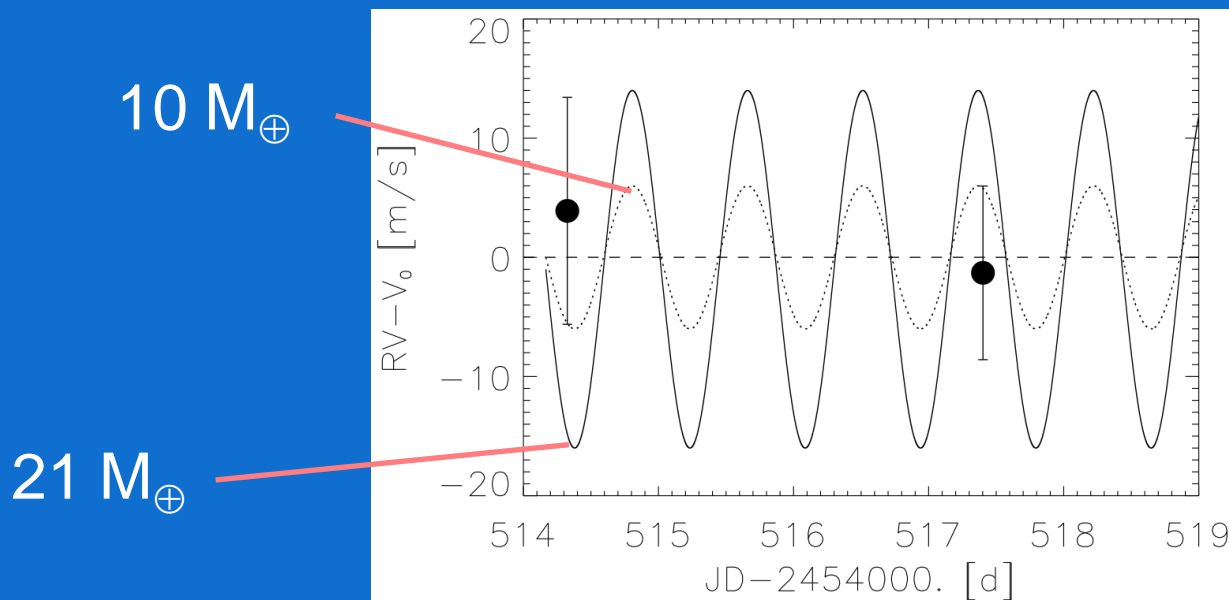
- If close binary ($P=0.85$ days) → Strong X-ray emitter
e.g. : YY Gem with Period = 0.81 days
- ROSAT satellite « all sky survey »
 $E = 0.1 - 2$ keV
region of Corot-Exo7b
- Results:
NO X-ray binary out to 250 ± 100 pc
- Late M binary system orbiting Corot-7b excluded





Follow-up 7: Radial Velocity

- Velocimetry with the spectrograph SOPHIE (Haute Provence)
 - 2 points measured
 - excludes a planet of mass $> 21 M_{\oplus}$ with 95% confidence level
- NO grazing Jupiter or white dwarf (excluded by timing anyway)





**All *known* cases of false positives
practically eliminated** with a high level
of confidence
(risk of a background blend = $8 \cdot 10^{-4}$)

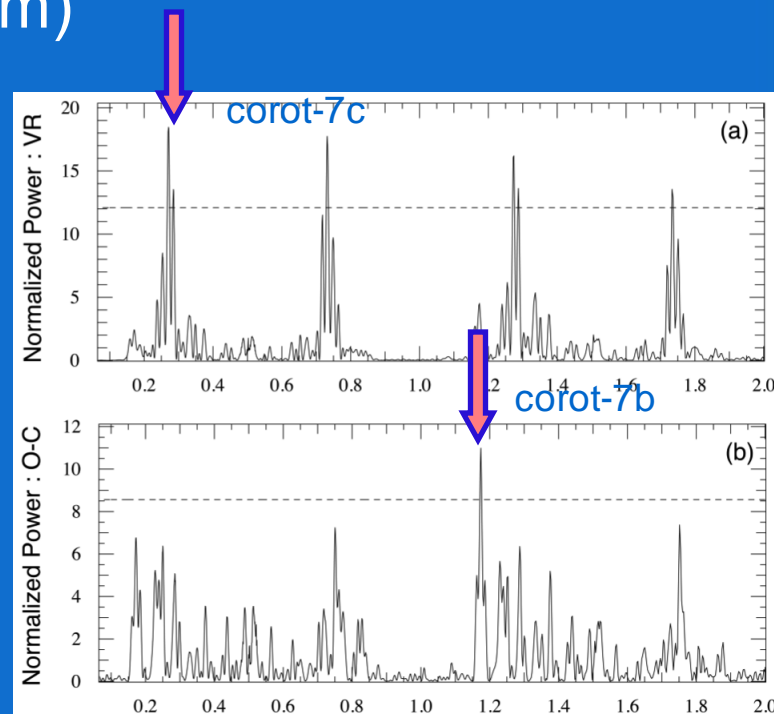
→ with a fairly high probability
the transit should be due to a
Super Earth planet
of radius $1.7 R_{\oplus}$
on a very close orbit

Léger, Rouan, Schneider et al., A&A, 2009



The real « Plus » : Radial Velocity with HARPS

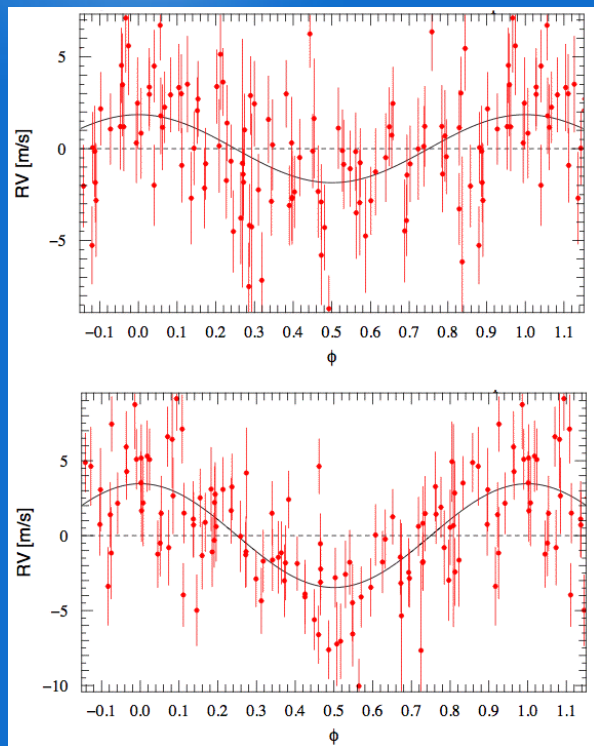
- HARPS Velocimetry (Queloz, et al. A&A 2009) :
 - 106 measurements (70h) during 4 months
- Difficulty : strong variability of the star affects the analysis
 - obvious correlation w rotation (23 days)
- Scargle analysis (\approx Fourier Transform)
 - after filtering of rotation (3 harmonics)
 - first peak : 3.7 days
 - Subtracted \rightarrow 2nd peak at 0.85 days
 - Phase of 0.85 d peak agrees w transit
 - Clear confirmation of Corot-7b
 - And discovery of *a second planet !*





Radial Velocity with HARPS

- Finally a *two Super-Earth* solution is found



Corot-7b
 $M = 4.8 \pm 0.8 M_{\oplus}$
 $P = 0.853$ days
 $a = 0.017$ AU
 $R = 1.68 \pm 0.1 R_{\oplus}$

Corot-7c
 $M = 8.4 \pm 0.9 M_{\oplus}$
 $P = 3.70$ days
 $a = 0.046$ AU

Parameters	Adopted solution
Period [days]	$0.853585 \pm 2.4 \cdot 10^{-5}$ (§)
T_{tr} [JD]	2454446.731 ± 0.003 (§)
T_0 [JD]	
K [m s ⁻¹]	3.5 ± 0.6
e	0
ω [deg]	180
m [M_{\oplus}]	4.8 ± 0.8
a [AU]	0.017
Period [days]	3.698 ± 0.003
T_0 [JD]	2454445.0 ± 0.7
K [m s ⁻¹]	4.0 ± 0.5
e	0
ω [deg]	180
m [M_{\oplus}]	8.4 ± 0.9
a [AU]	0.046

- First case of a system with 2 Super-Earths, one in transit
- A third planet is suspected → a very compact system !



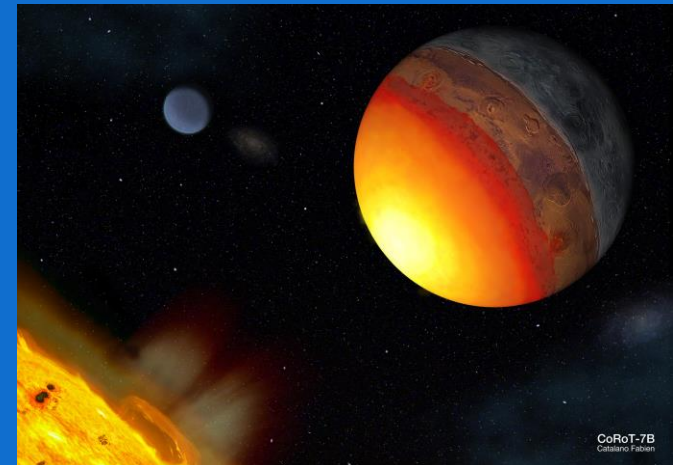
Corot-7b characteristics

- A big sister of the Earth :
 - $M = 4.8 - 6 M_{\oplus}$
 - $R = 1.6 - 1.7 R_{\oplus}$
 - Deserves the Super-Earth denomination
- A very different *ecologic* environment
 - Period: $P = 0.85 \text{ d}$ → one year = 20.5 h !
 - Orbit radius : $a = 0.017 \text{ AU} = 4.5 \text{ stellar radius !}$
→ A sun of 28° in the sky !



Nature of Corot-7b ?

- First solid evidence for a rocky planet !
 - A lot of exciting physics (Léger et al. Submitted)
- Tidal forces :
 - Must be **phase-locked** ($t_{\text{synch}} < 100$ yrs) :
keep one face towards the star (as the Moon vs Earth)
 - Moderately elongated under tidal forces < 150 km
 - Tidal heating could be extremely efficient if *excentricity* is $\neq 0$
could be *analog to Io* → *volcanism* ?

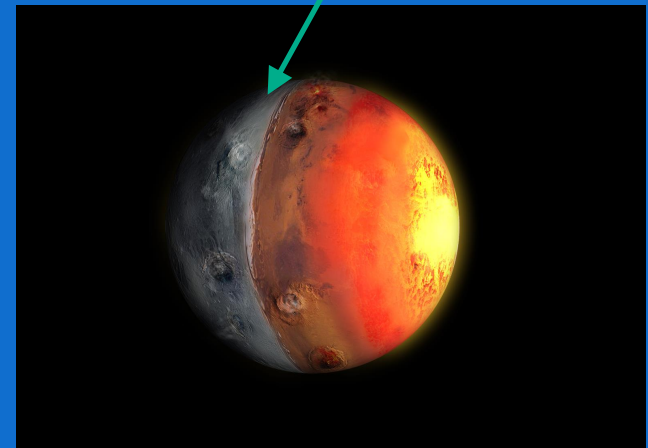
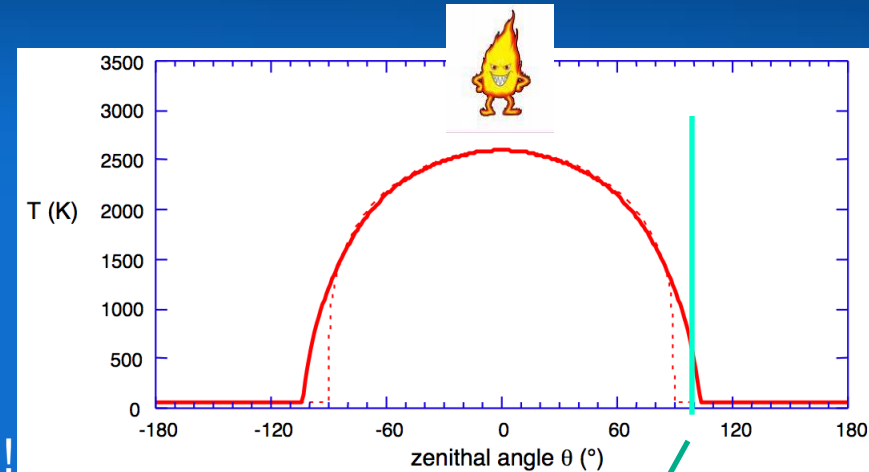




Nature of Corot-7b ?

- Temperature

- dark side : 50K pretty cold !
geothermal origin
- dayside : 1800-2600 K : pretty hot !
no heat redistribution by atmosphere or ground
- Thus... a *temperate* region of 90 km width only (270-370 K)
 - life is however not possible : no water !

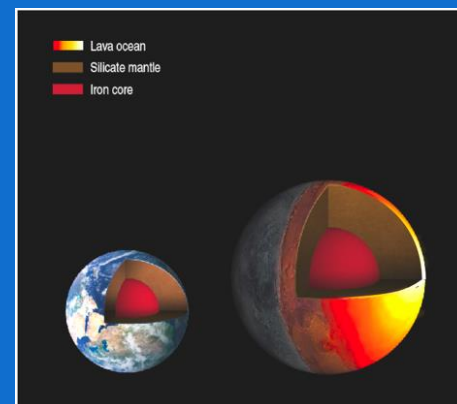
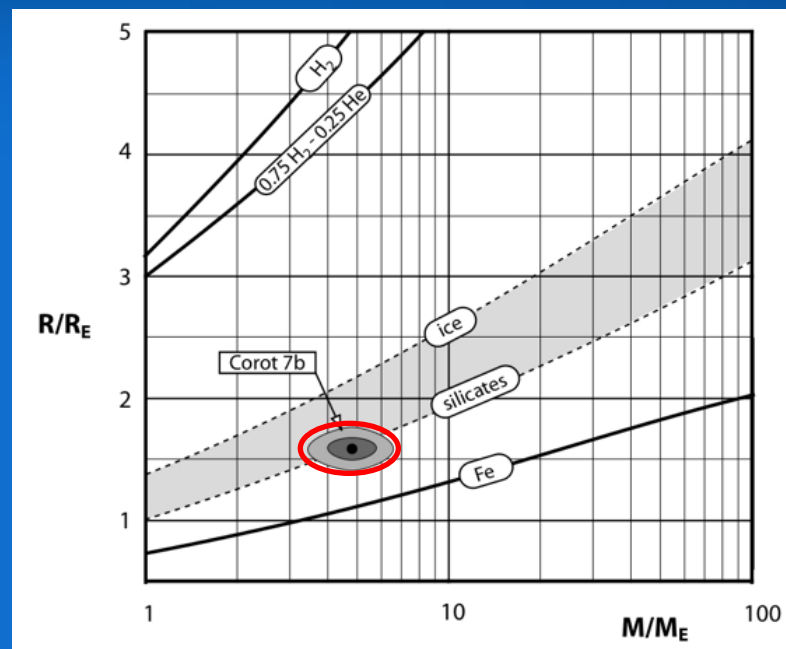




Nature of Corot-7b ?

● Structure

- First time density of a SE measured
- $5.5 - 7 \text{ g.cm}^{-3}$ (~Earth) → rocky
Most probably silicates mantle compatible w 20% of water
- Likely migrated from a wider orbit
core of a Neptune
or a true rocky telluric
- surface : *ocean* of liquid rock
up to a latitude of $\approx 40^\circ$
refractory oxides: CaO , Al_2O_3
- continent or frost H_2O , CO_2 elsewhere

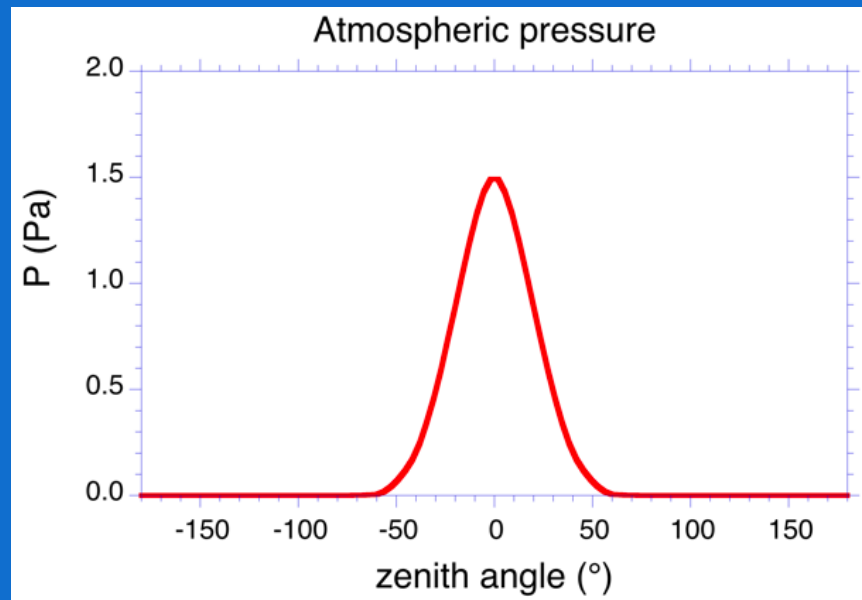




Nature of Corot-7b ?

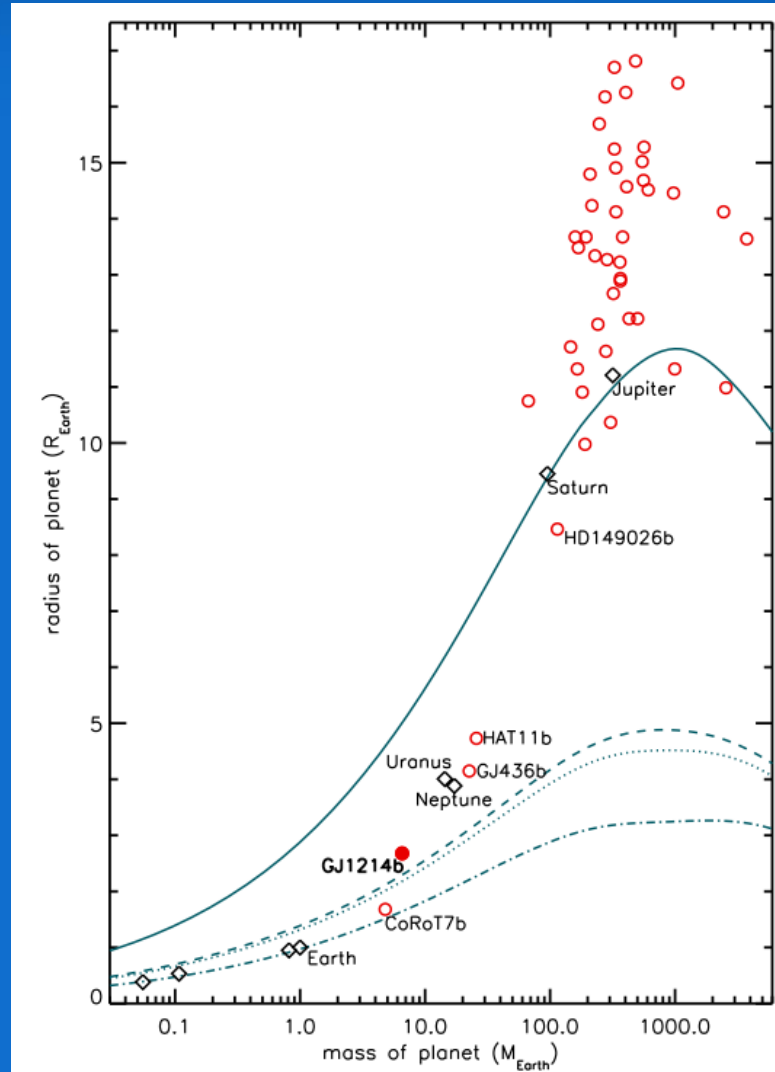
● Atmosphere

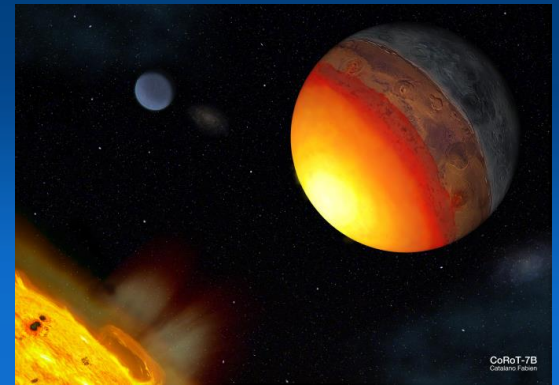
- Escape of all volatiles within 0.1 to 1 Gyr (< age of the system)
- Composition : vaporised silicates
 - Should be extremely tenuous : $P = 5-0.05 \cdot 10^{-5}$ bar





Telluric or icy giant ?





Corot-7b :

**A planet whose mass and size make
it a cousin of Earth**

Orbiting a star cousin of our Sun

**→ A solid milestone on the pathway
towards habitable planets**

Thanks extended to both CoRoT and HARPS teams



Thank you for your attention

